# No.7

# 陰影解析

# Shape from intensity

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# Shape-from-X

#### Shape estimation from "X" as a clue



Stereo / Motion



Triangulation



Texture



Focus / Defocus



Time

### Shape-from-Intensity

Shape estimation with observed intensity as a clue

- Inverse process of rendering to determine pixel intensity
  - Inverse geometry

Illuminations and reflectance properties
Unknown: Scene shape

Clue

■Shading, specular reflection, shadow, etc.



# Shape-from-Shading

### Shape estimation based on shading

- Photometric information contains a lot of clues about geometry.
  - Surface normalDepth
- Start with a simple problem setting
   A point light source at infinity (parallel light)
   Fixed camera and target object
   Perfect Lambert reflection

(No specular reflection, No shadow)■No global illumination



### Surface normal



Partial differential of z with x and y to define the inclination.

$$-\frac{\partial z}{\partial x} = \frac{A}{C} = p \qquad -\frac{\partial z}{\partial y} = \frac{B}{C} = q$$

Unit normal vector from the inclination.

defined by two parameters

$$\mathbf{n} = \frac{\left(p, q, 1\right)^{\mathrm{T}}}{\sqrt{p^2 + q^2 + 1}}$$

# Illumination and reflection

### Parallel illumination

Ideal illumination coming from a point light source at infinity
 Every surface point is illuminated with the same irradiance from the same direction

Unit parallel illumination vector

$$\mathbf{s} = \frac{(p_s, q_s, 1)^{T}}{\sqrt{p_s^2 + q_s^2 + 1}}$$

**n** normal

#### Lambert diffuse reflection

Ideal reflection that uniformly reflects incident light in all directions

 $(\rho \ge 0)$ 



 $\square \rho$  : Lambert diffuse reflectance

$$i = \rho \mathbf{s}^{\mathrm{T}} \mathbf{n}$$

Problem setting

# Known: illumination and observed intensitiesUnknown: normal

Unknown 
$$\mathbf{n} = \frac{(p,q,1)^{\mathrm{T}}}{\sqrt{p^{2}+q^{2}+1}}$$
  $\mathbf{s} = \frac{(p_{s},q_{s},1)^{\mathrm{T}}}{\sqrt{p_{s}^{2}+q_{s}^{2}+1}}$  Known

For simplicity, assuming that diffuse reflectance  $\rho = 1$ 

$$I = \cos \theta_i = \mathbf{n}^{\mathrm{T}} \mathbf{s} = \frac{\left(pp_s + qq_s + 1\right)}{\sqrt{p^2 + q^2 + 1}\sqrt{p_s^2 + q_s^2 + 1}} = \frac{R(p,q)}{\operatorname{Reflectance map}}$$



# Shape-from-Shading

The normal (*p*, *q*) cannot be uniquely determined from one observed intensity

One equation and two unknown parameters

Need to add some information

Assumption that the object surface is smooth

Prior knowledge of shape

Increase Illumination directions



# Photometric Stereo(照度差ステレオ)

- Taking multiple images with changing illumination directions
- Solving ambiguity in the reflectance map
- Assuming Lambert diffuse reflection, it can be solved linearly



# Example of Photometric Stereo



Normal map

Input images

In the case of two light sources

The intersection of the two solution curves becomes new solution space



### In the case of three light sources



Linear solution in the case of three light sources

Assuming that three observation intensities  $(i_1, i_2, i_3)$ were obtained for a pixel under three different illumination directions  $(s_1, s_2, s_3)$ 



Stable because it can be solved linearly.
 Reflectance (*ρ*) is also estimated at the same time.

# In the case of more than three light sources

Assumed that observed brightness  $(i_1, i_2, ..., i_M)$  was obtained under M (> 3) different illumination directions  $(s_1, s_2, ..., s_M)$ 



Since the illumination matrix S is not a square matrix, calculated using a pseudo inverse matrix

$$\widetilde{n} =$$
mini-report

Moore Penrose(ムーア・ペンローズ)inverse matrix

Least squares method assuming that observation error is Gaussian distribution

more stable and accurate solution

# The merit of multiple light sources

Avoid specular reflection and shadow
 Assuming that pure Lambert diffuse reflection can be observed in at least three images.



without

with



Estimated reflectance

Shading generated from the estimated normal

Barsky, S, et al, The 4-source photometric stereo technique for three-dimensional surfaces in the presence of highlights and shadows

# Photometric Stereo in Parthenon



Near light source photometric stereo.

Two black hemispheres to determine light source positions.



[Per Einarsson et al., Photometric Stereo for Archeological Inscriptions, 2004]

# Summary of the number of light sources

- 1 light sourceShape-from-Shading
- ■2 light sources ■Photometric Stereo (照度差ステレオ法)
- 3 light sources
  - Uniquely solved Photometric Stereo
  - Simultaneous estimation of reflectance and normal
- More light sources
  - Robust to specular reflection and shadow

# 3D shape and normal

Even if normals are known, 3D shape cannot be uniquely determined.

Height ambiguity due to different integral path.Differential is easy, but integration is difficult.



# Integrability (可積分性)

When integrated along a closed loop on a smooth surface, the integral value becomes zero.

□Independent to the integral path.

 $\square Minimization of objective function E(Z)$ 



Estimation of light source direction/position

Use reference object with known shapeEstimate light source direction by a black sphere

Estimate near light source position by two or more black spheres

- In the outdoor, high dynamic range measurement is required
  - Mirror sphere for ambient light
  - Black sphere for position of sun
  - Diffuse sphere for brightness of ambient light



# When the illumination direction is unknown

■Illumination directions and strengths are unknown: □Uncalibrated Photometric Stereo(未校正照度差ステレオ)

$$\begin{bmatrix} i_{11} & \cdots & i_{1N} \\ \vdots & \ddots & \vdots \\ \vdots & \ddots & \vdots \\ i_{M1} & \cdots & i_{MN} \end{bmatrix} = \begin{bmatrix} \mathbf{s}_1^{\mathrm{T}} \\ \vdots \\ \mathbf{s}_M^{\mathrm{T}} \end{bmatrix} [\rho_1 \mathbf{n}_1 & \cdots & \rho_N \mathbf{n}_N]$$

Singular value decomposition  $\mathbf{I} = \mathbf{S} \widetilde{\mathbf{N}} = \mathbf{U} \mathbf{\Sigma} V^{\mathrm{T}}$ 

(特異値分解)  

$$\begin{cases}
\mathbf{S} = \mathbf{U}'(\mathbf{\Sigma}')^{\frac{1}{2}} & \mathbf{I} & \mathbf{U}' \mathbf{\Sigma}' & \mathbf{V}^{T'} \\
\mathbf{I} & \mathbf{U} & \mathbf{\Sigma} \\
\mathbf{I} & \mathbf{U} & \mathbf{\Sigma} \\
\mathbf{I} & \mathbf{I} & \mathbf{I} & \mathbf{I} \\
\mathbf{I} \\
\mathbf{I} & \mathbf{I} \\
\mathbf{I} &$$

# Bas-Relief Ambiguity (浅浮き彫りの曖昧性)

Decomposition of a matrix is not uniquely determined.

 $\mathbf{I} = \mathbf{S}\widetilde{\mathbf{N}} = (\mathbf{S}'\mathbf{H})(\mathbf{H}^{-1}\widetilde{\mathbf{N}}')$ 

 ${\bf H}$  is any 3  $\times$  3 matrix

Linear uncertainty
 Different incident light and normal pairs produce the same shading





"The Bas-Relief Ambiguity", Belhumeur et al, 1997

# Shape-from-Shadow

Limited existing space in the cone
 top vertex: light source
 bottom surface: shadow region
 Logical AND of many cones
 Estimate rough convex hull shape (凸包形状)



Light sources cone Shadow region

# Shadow graphs

#### Restrict the existing space of objects from shadows







A part of the 48 input images



Shape estimation [Shadow Graphs, Yu and Chang 2005]

# Shape-from-Specularity

# Shape estimation based on specular reflection Strong specular reflections around the mirror direction



A part of 65 input images





Camera (12bit, 1300x1030)

Metallic ball for estimating light source position



[Chen et al., Mesostructure from Specularity (CVPR2006)]

# Scatter-Trace Photography

- Shape estimation based on specular reflection
- Complex transparent scene with inhomogeneous interior
   Observed reference pattern on the transparent object
   Shape estimation by hypothesis and verification



Example of Scatter-Trace at each point on the surface

Nigel Morris and Kiriakos Kutulakos,

Reconstructing the Surface of Inhomogeneous Transparent Scenes by Scatter-Trace Photography, ICCV2007.

# Analysis of Scatter-Trace

If direct reflection on the surface, Scatter-Trace stripes
 intersect with the eye and the surface of the object
 monotonically decrease as the distance increases

 $T_{\mathbf{a}}^{\mathbf{I}}$ 







depth normal



#### Verification,

Removing non-monotonically decreasing component → If the direct reflection components remain, the assumed depth is correct.



# Example-based Photometric Stereo

[Aaron Hertzmann and Steven M Seitz, 2005]

#### How to handle any BRDF?

Reference object whose shape is known and BRDF is same with the target object to be measured.

Used as a look up table to find similar reflection properties.





Examples of the input images

Estimated normal

If reference objects can be prepared, any BRDF can be treated.





# Separation of reflection components (反射成分の分解)

### Real scene

Many phenomena mix in real image.

Use complex model or decompose in advance.





Again

### Dichromatic reflection model (Shafer 1985) (2色性反射モデル)

#### Reflected light = Diffuse reflection + Specular reflection

- ■Diffuse reflection (拡散反射): Reflection inside the surface layer Object color
- ■Specular reflection(鏡面反射): Reflection at the border between air and surface layer □Light color









Specular reflection Sum of both reflection

Diffuse reflection

# Dichromatic reflection model

#### Decomposition based of color difference



# Viewpoint dependency

Diffuse reflection :

Specular reflection :

n: dependent or independent or independent? On viewpoint

Assume uniform specular reflection characteristics
 Simultaneous estimation of light source distribution from specular reflection components

dependent or









Input image (different viewpoint ) Diffuse reflection Specular reflection

# Difference in polarization (偏光)

- Separate specular reflection components using polarization
- Simultaneous estimation of specular reflection parameter and near light source position
- Assuming monochromatic uniform reflection characteristics





Input image Diffuse reflection component Specular reflection component

## RANSAC-based method



Convert real image into pure Lambert image based on RANSAC (RANdom SAmple Consensus)

Choose three images randomly from the input image set, converted to fully satisfy Lambert model (linearization)

Remove shadows and specular reflections as outliers



conversion

```
\mathbf{i} = \mathbf{S}^{\mathrm{T}} \mathbf{\widetilde{n}}
```

Specular reflection



One of the input image

Linearized image (Negative: red)

Estimation of optical phenomena

Y. Mukaigawa, et al, ``Analysis of photometric factors based on photometric linearization'', JOSA2007

Some approaches to handle more complex scene

Preprocess to extract pure Lambert diffuse reflection component

Optically or mathematically

Use mode complex modelSince the parameter increases, it may become unstable

Solve in the framework of robust estimation

Use many input images and consider non-Lambert components as outliers

Solve by deep learning



### Summary

From the shading information, not only photometric information such as color and reflection properties, but also geometric information such as normal can be extracted.

Photometric stereo cannot estimate <sup>mini-</sup><sub>report</sub>. It can estimate mini-report

Many traditional methods assume Lambert diffuse reflection, but extended to more complex scene in recent years.